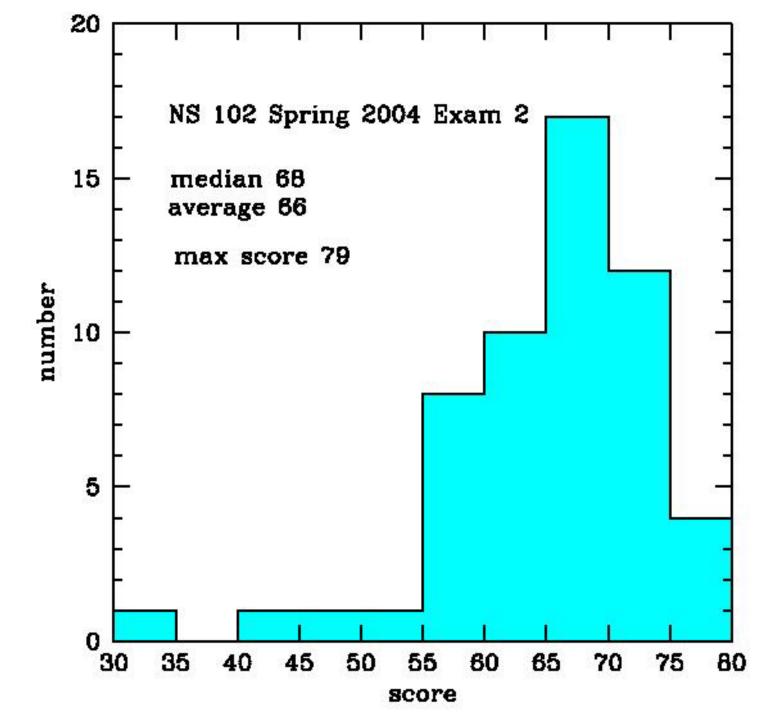
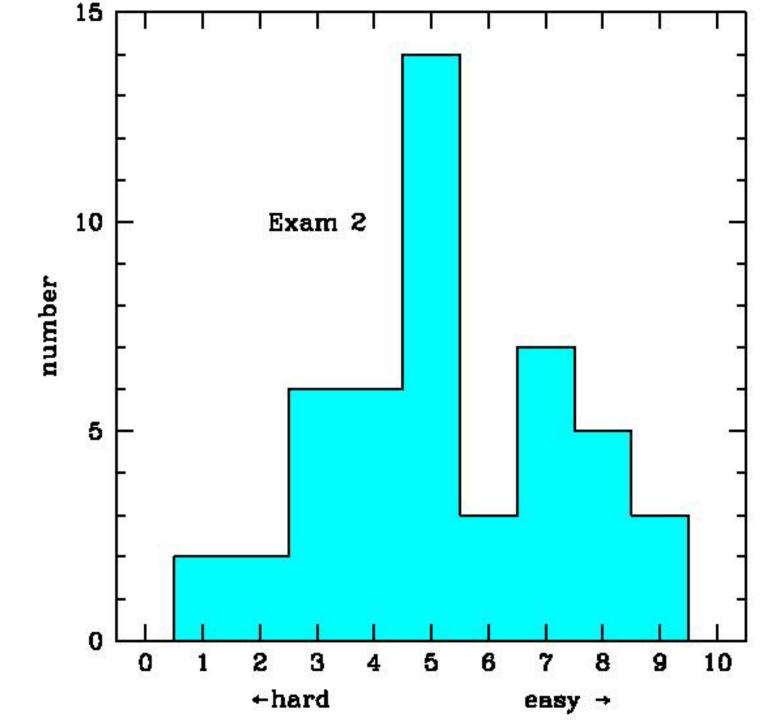
## Lecture 15 May 25, 2003

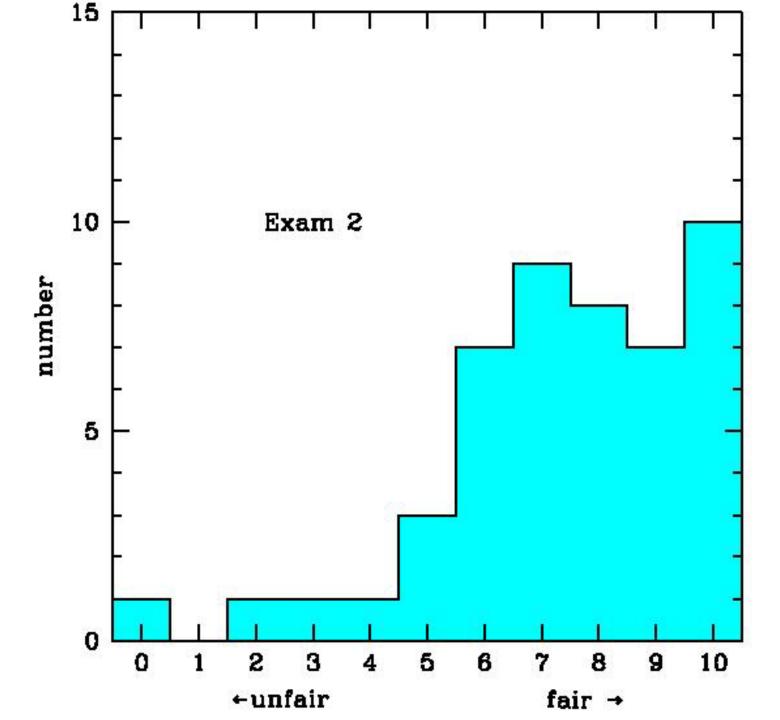
## What is the Universe made of?

## News of the week

- Regular office hours Tuesday & Thursday (next week also)
- Lab this week: The mass density of the Universe
- Finish reading Hawking
- Lectures:
  - May 25<sup>th</sup>: What is the Universe made of?
  - May 27<sup>th</sup>: The first second.
  - June 1<sup>st</sup>: It's about time.
- Final Exam: Tuesday, June 8th, 10:30am-12:30pm
- Sarah Stein (graduating student) please see me.





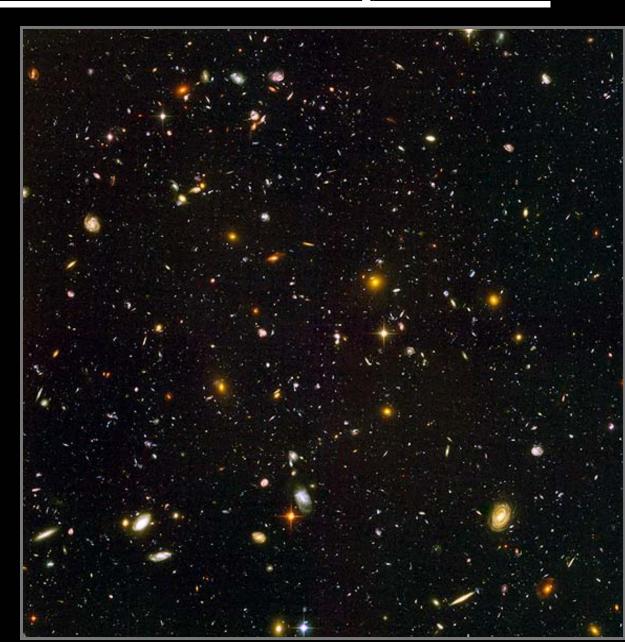


#### The Hubble Ultra Deep Field

UNIVERSE OF GALAXIES

3000 here

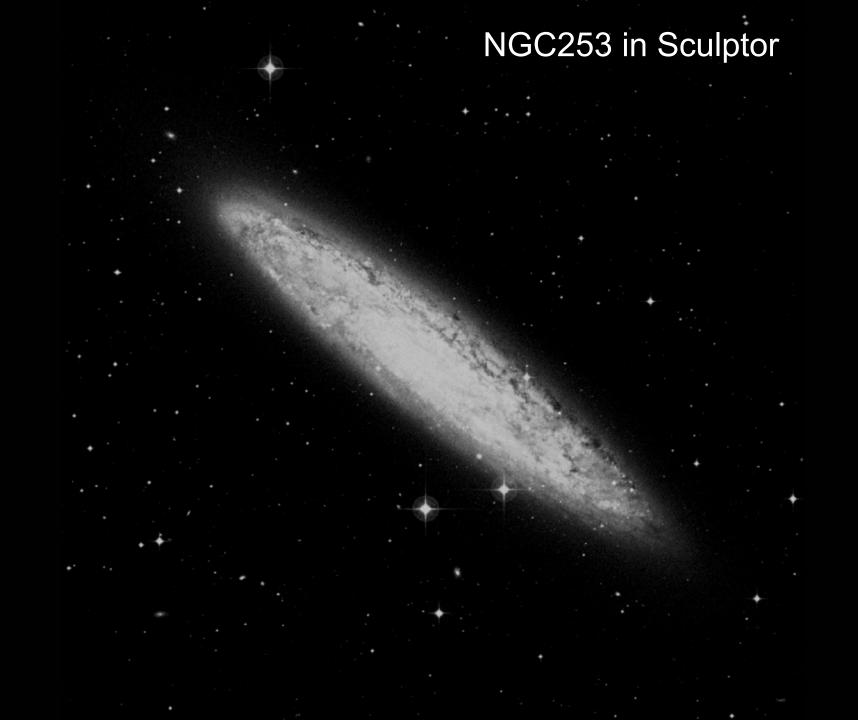
50 billion over entire sky



## The visible universe

M63

Galaxies



$$R \rightarrow R$$

$$\frac{\mathbf{v}^2}{R} = \frac{GM_{SUN}}{R^2}$$

measure v & R

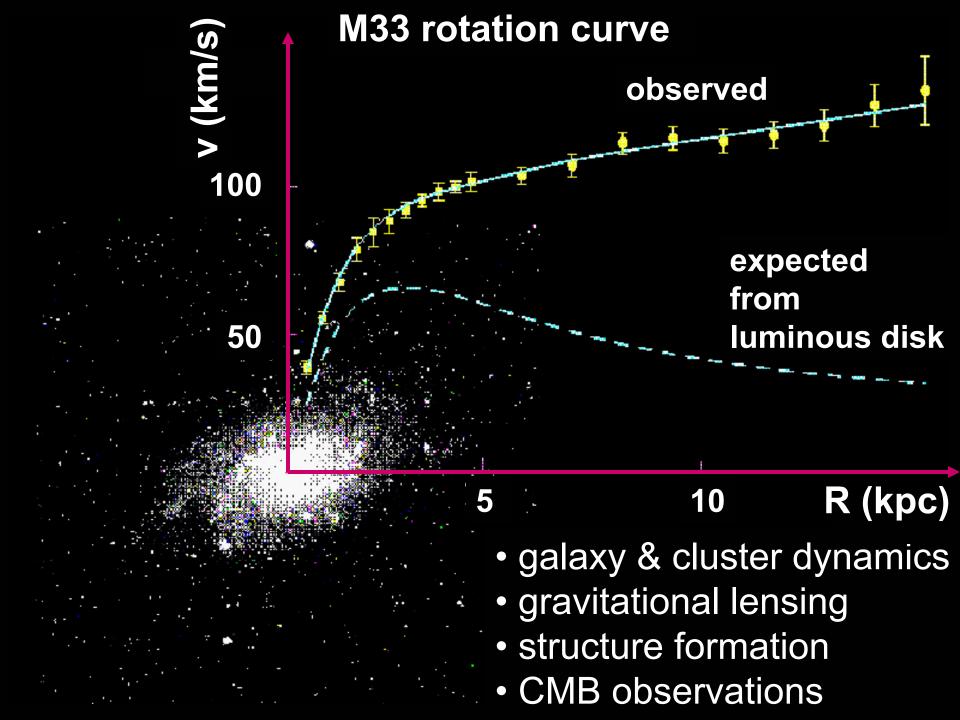


 $\longrightarrow M_{SUN}$ 



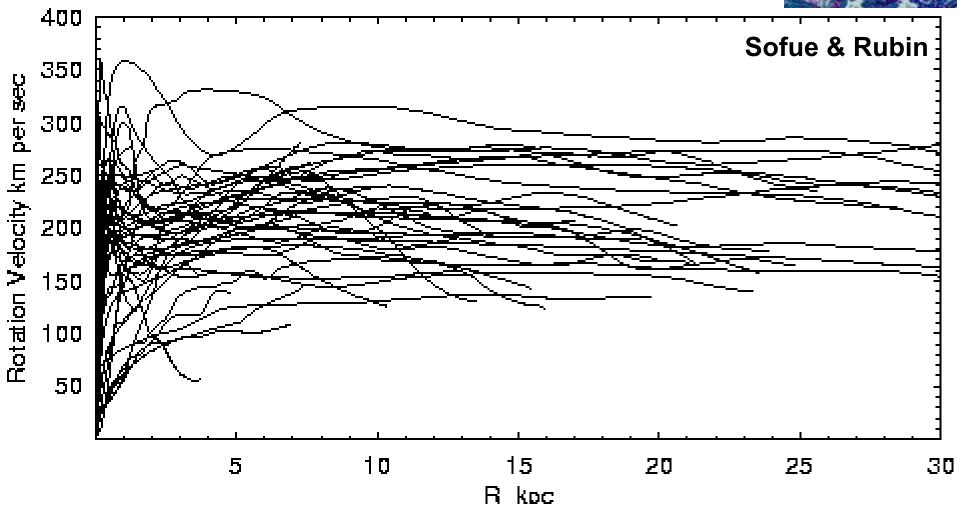
measure v & R

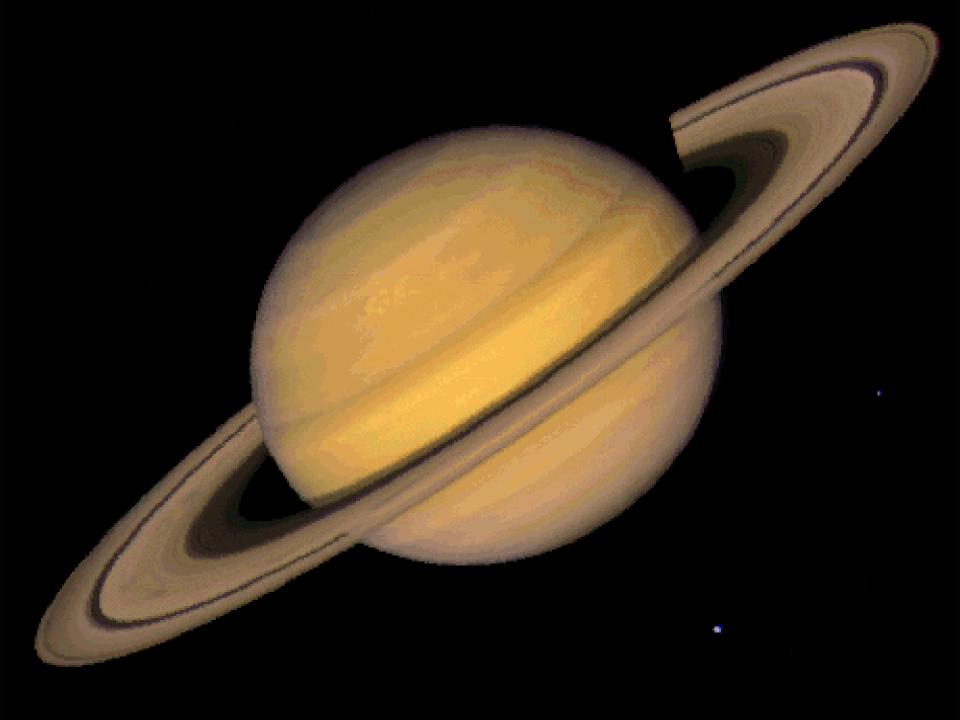


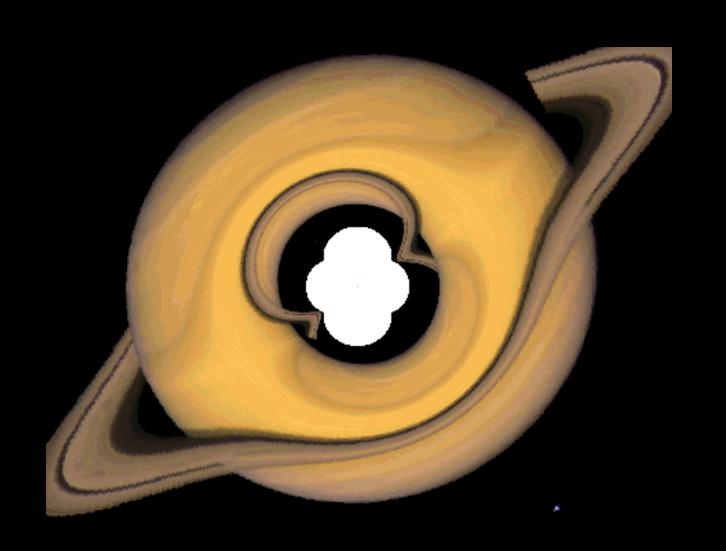


## Rotation curves











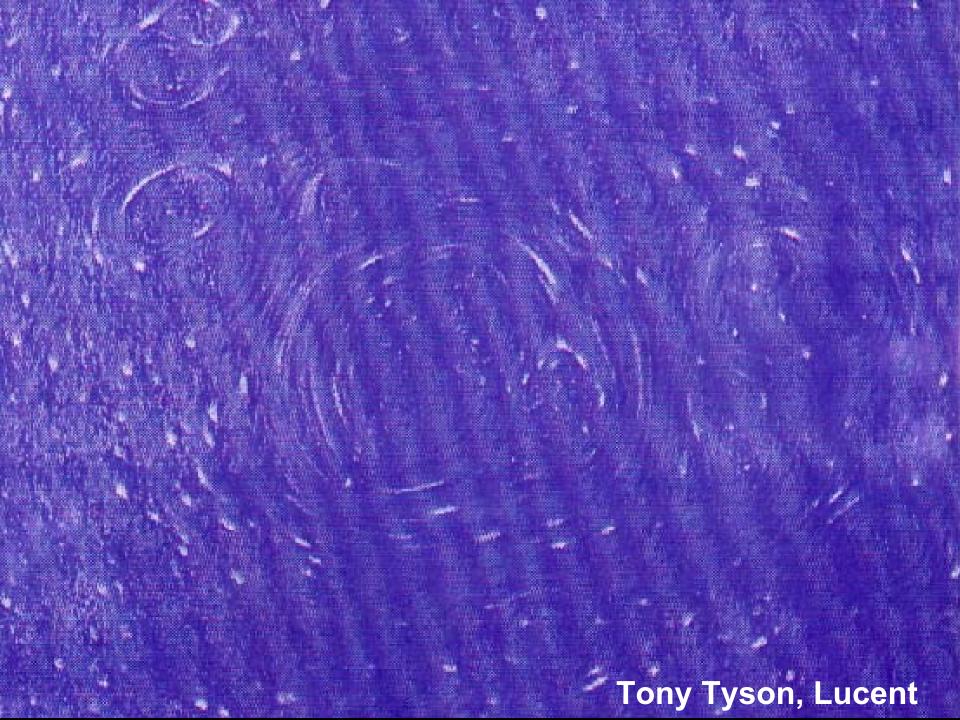
Gravitational Lens in Abell 2218
PF95-14 · ST Scl OPO · April 5, 1995 · W. Couch (UNSW), NASA

HST · WFPC2



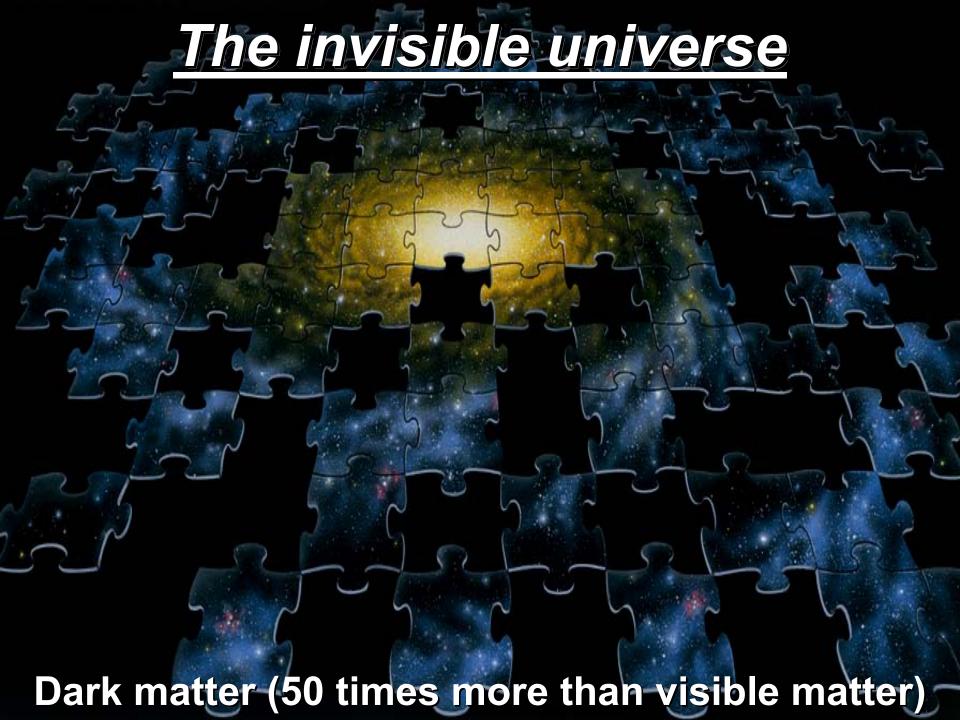
## Gravitational Lens Galaxy Cluster 0024+1654

Hubble Space Telescope • WFPC2

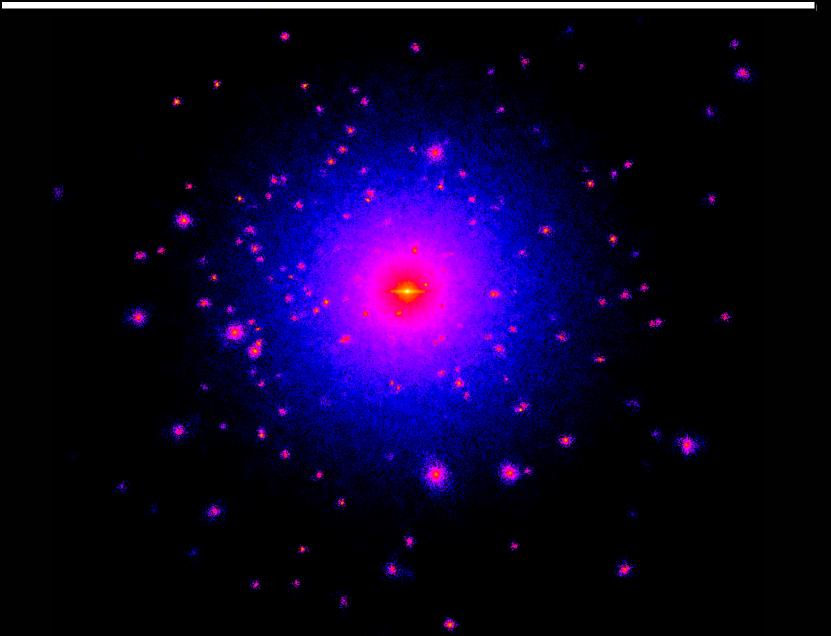




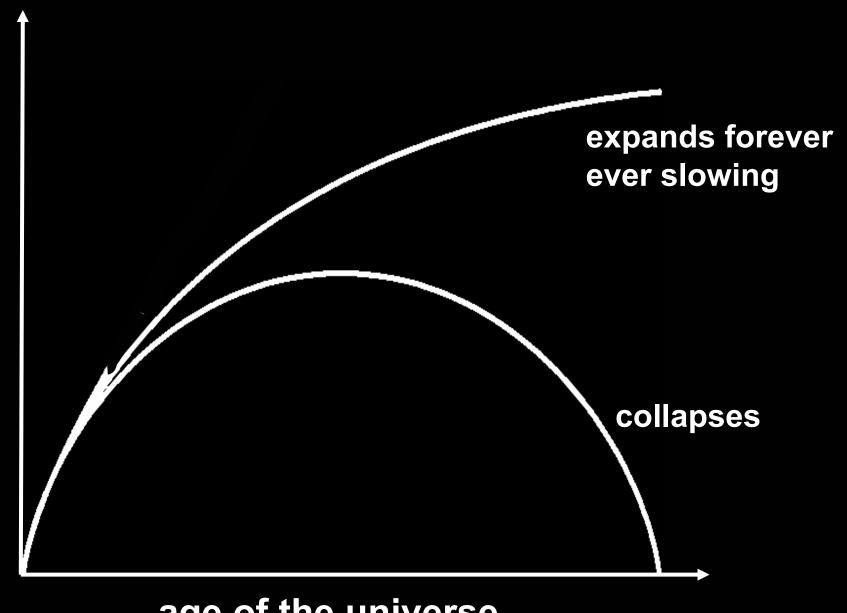
#### Most of the universe is dark!



# If we could "see" dark matter



## The critical density



age of the universe

distance

## The cosmic food chain $(\Omega_i)$

The critical density:

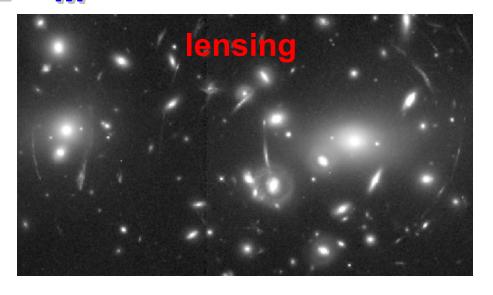
$$\rho_C = \frac{3H_0^2}{8\pi G} \simeq 10^{-30} \text{ g cm}^{-3}$$

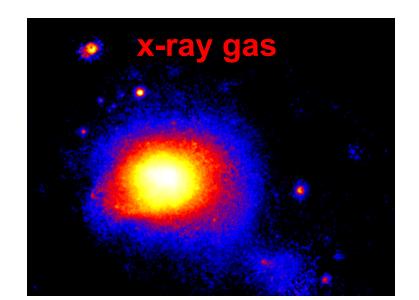
Fraction of the critical density:

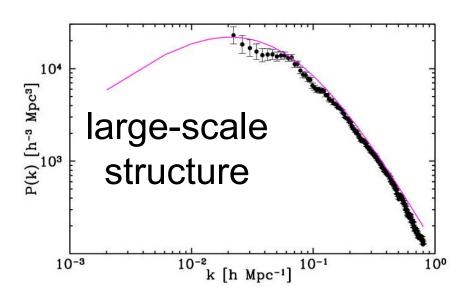
$$\Omega_i = \frac{\rho_i}{\rho_C}$$

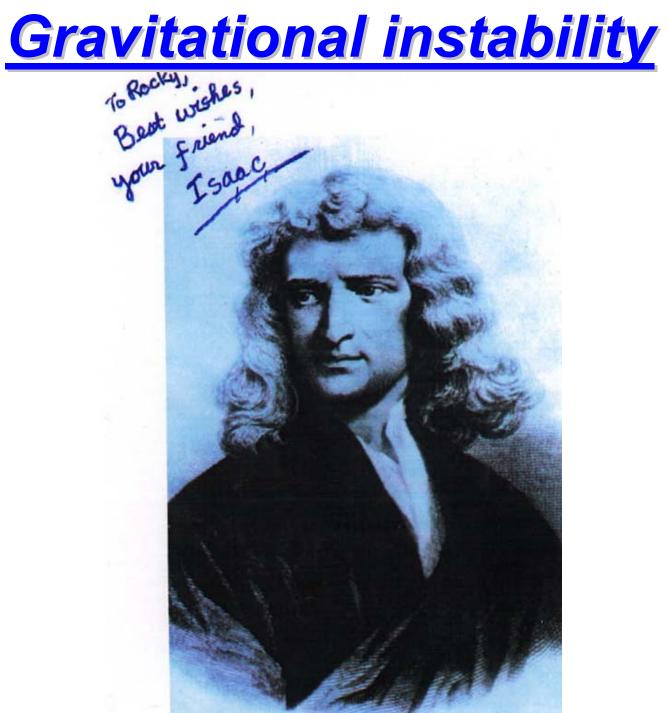
## <u>Matter</u> $\Omega_{\rm M} \sim 0.3$

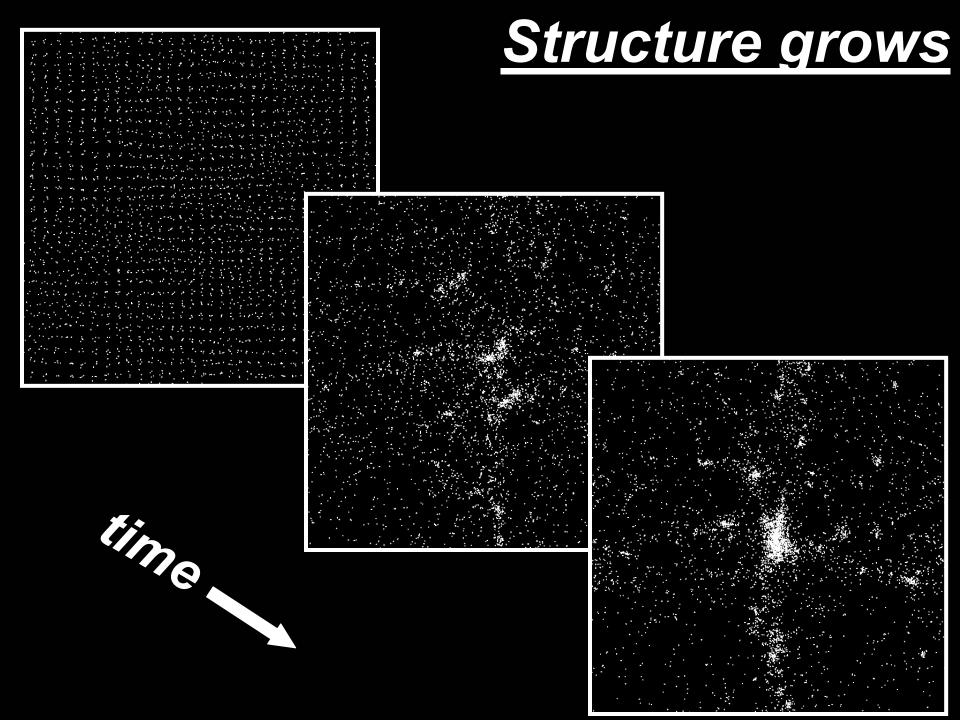






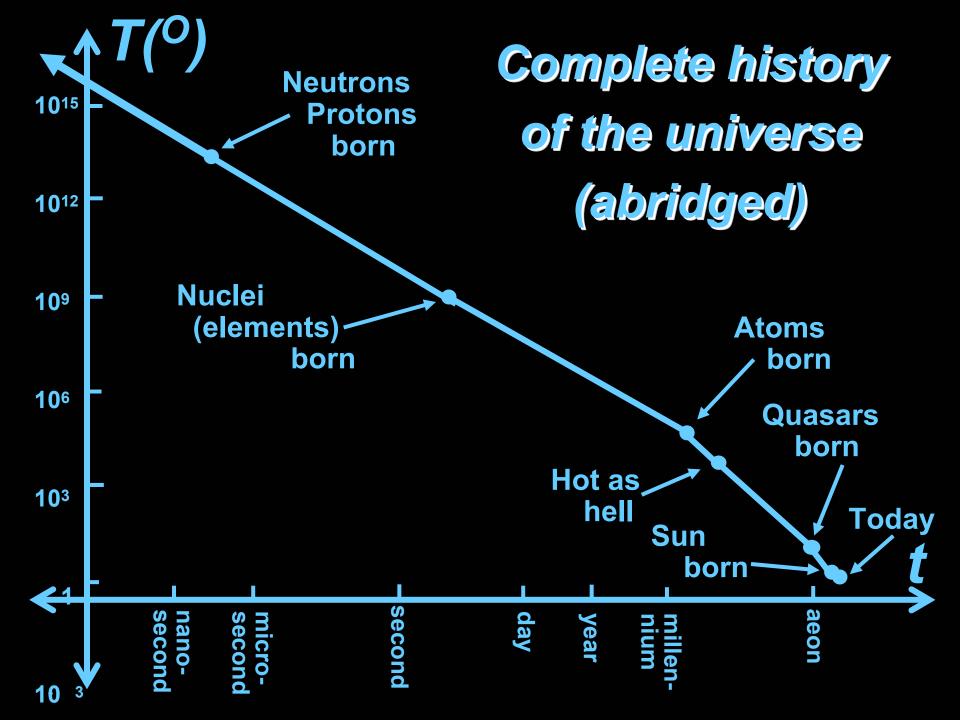






#### Most of the universe is dark!

It ain't even normal stuff!

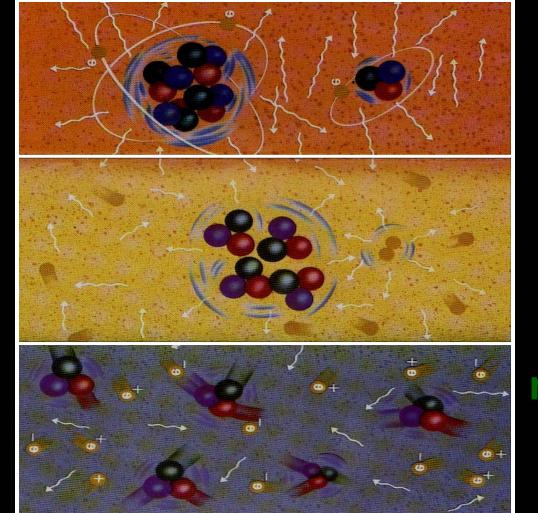


380,000 years

3 minutes

1-micro second

4-pico seconds



atoms form

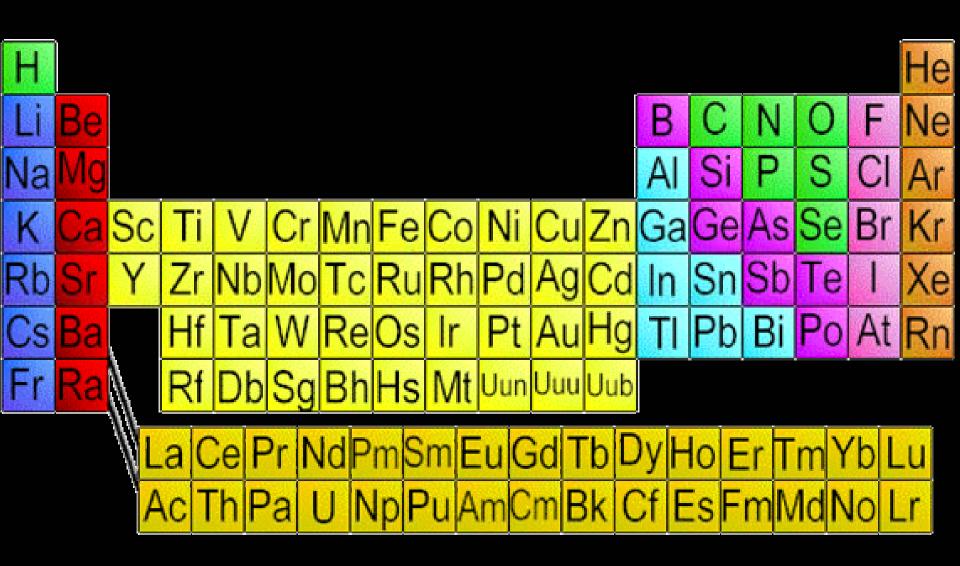
nuclei form

neutrons protons form

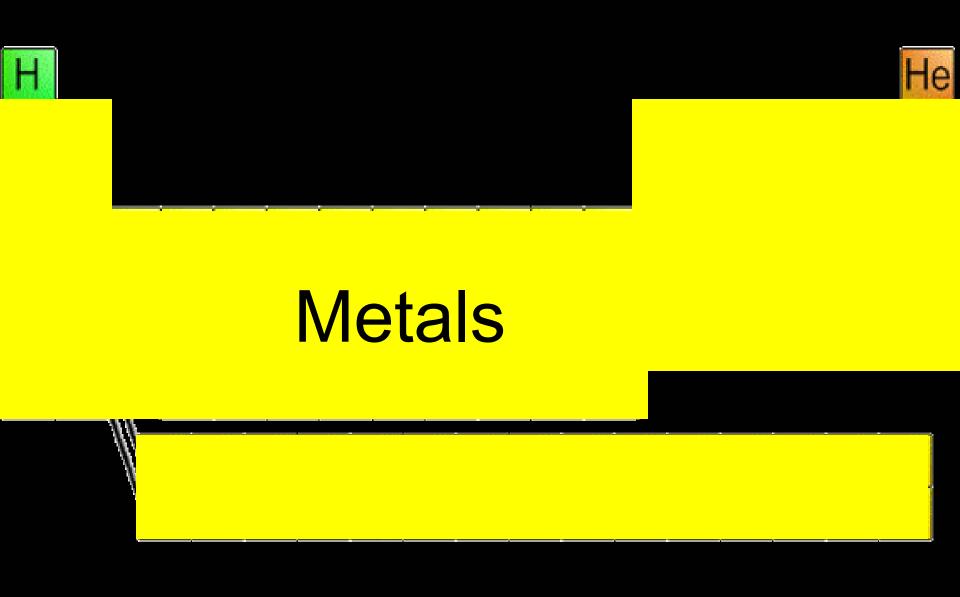
primordial soup

**BANG!** 

#### Periodic table - chemist



## Periodic table - cosmologist



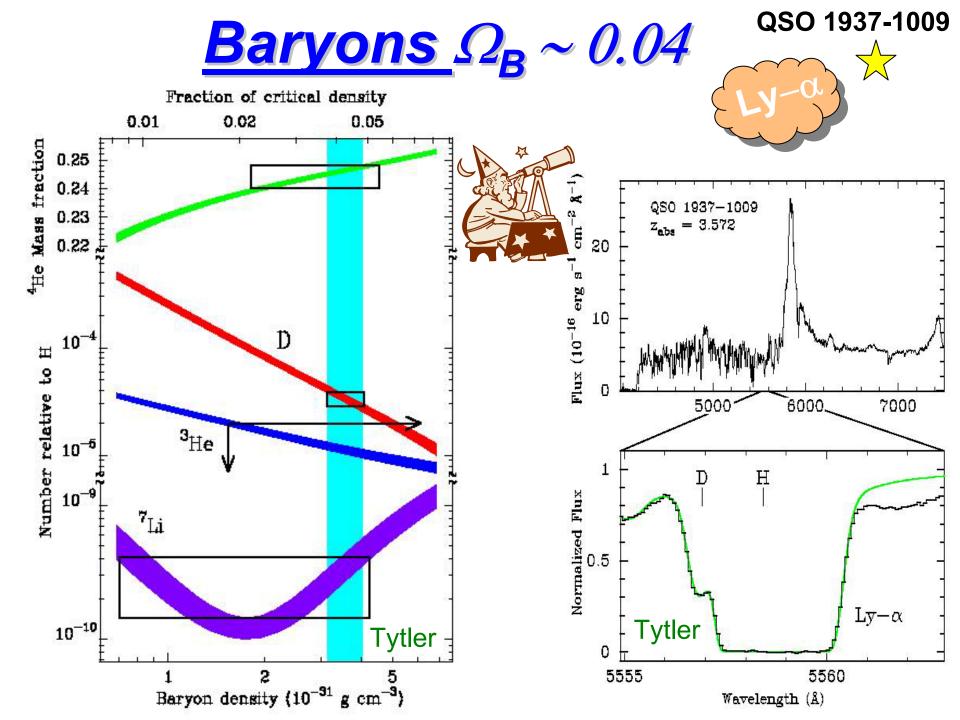
## The Chemical Elements

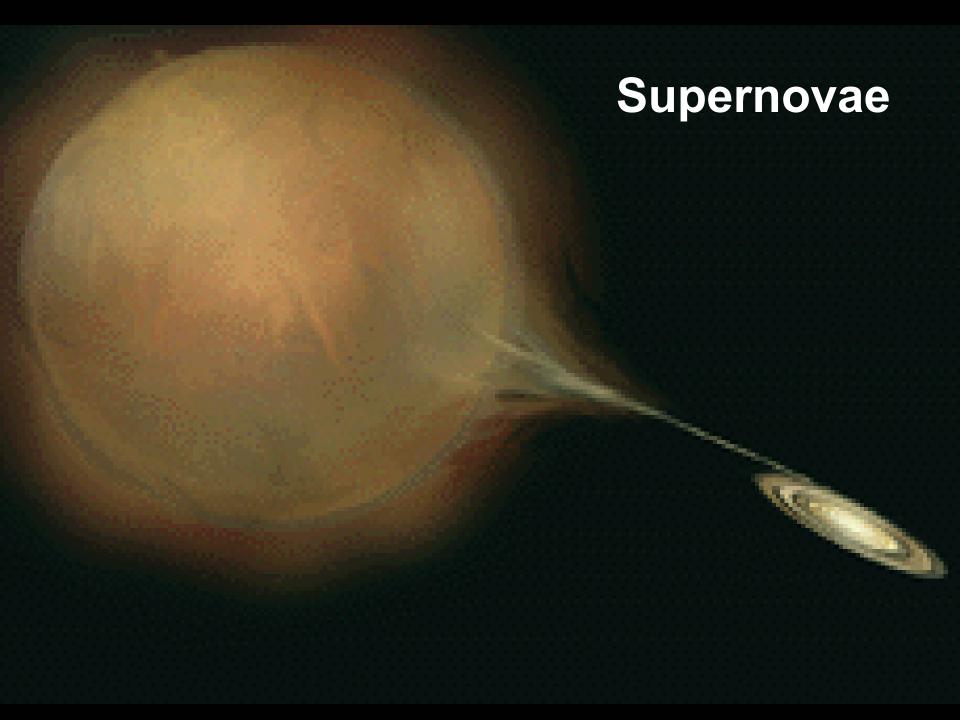
## Today:

```
73% Hydrogen (10<sup>-5</sup> deuterium)
26% Helium (10<sup>-5</sup> <sup>3</sup>He)
1% Metals
```

#### Three minutes AB:

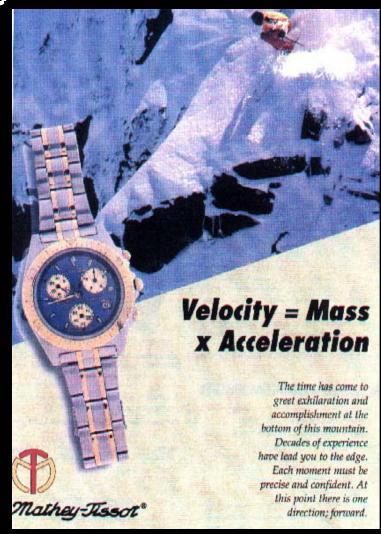
```
76% Hydrogen (10<sup>-5</sup> deuterium)
24% Helium (10<sup>-5</sup> <sup>3</sup>He)
0.0000001% Lithium
```





# Dark matter?

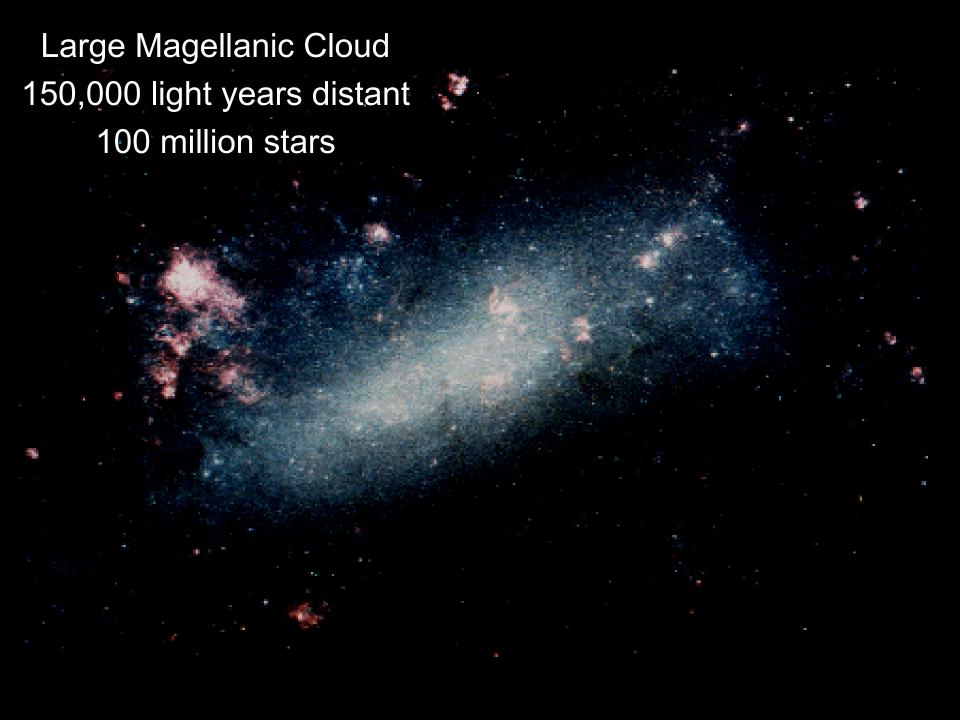
Modified Newtonian dynamics

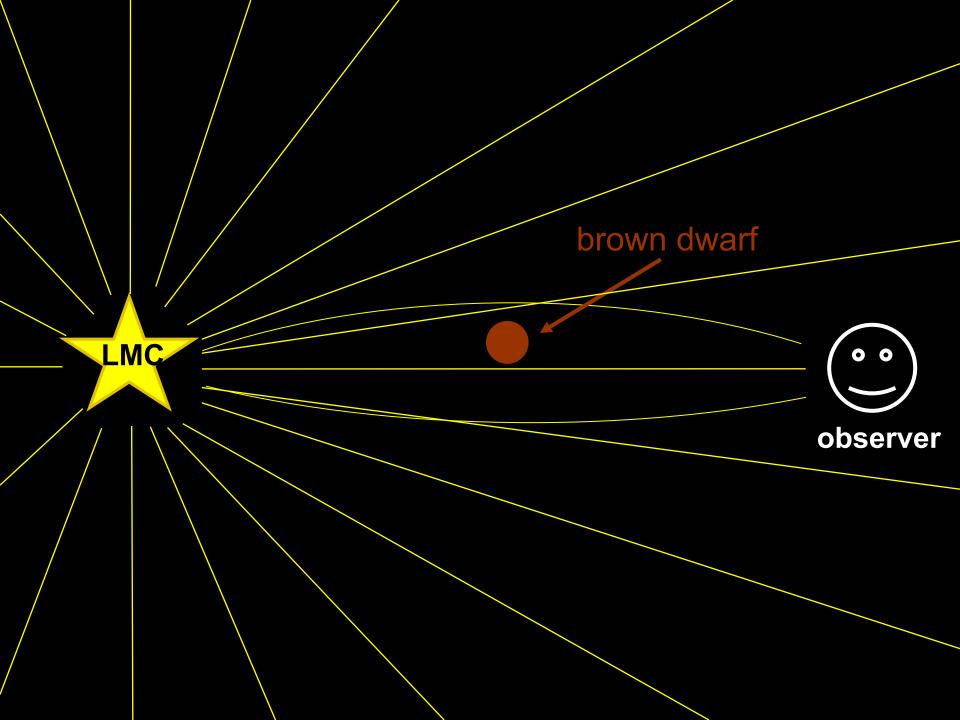


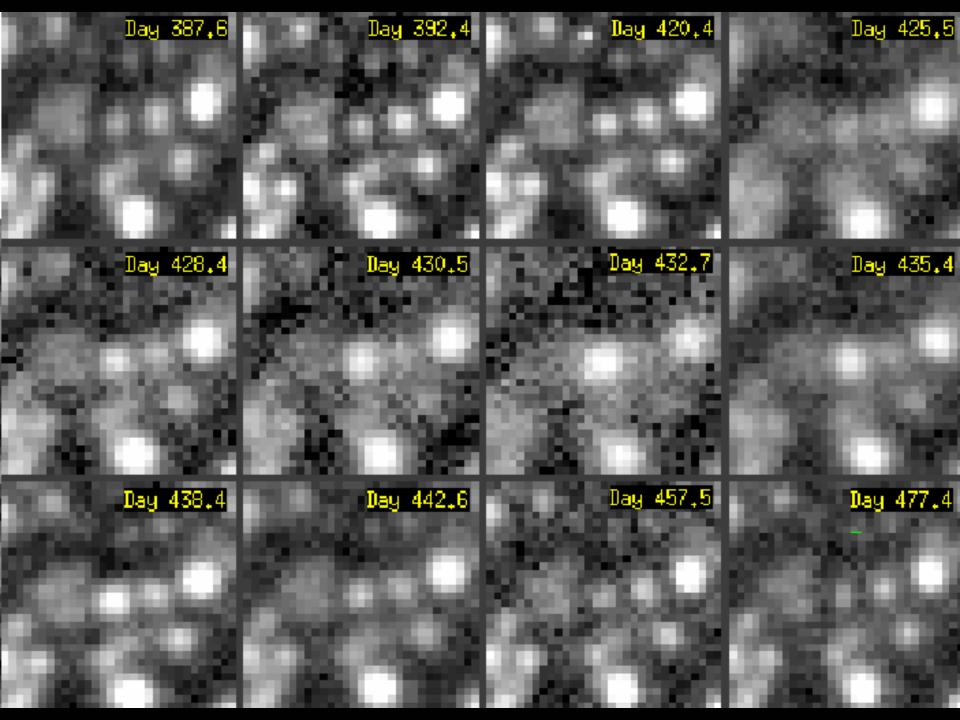
# Dark matter?

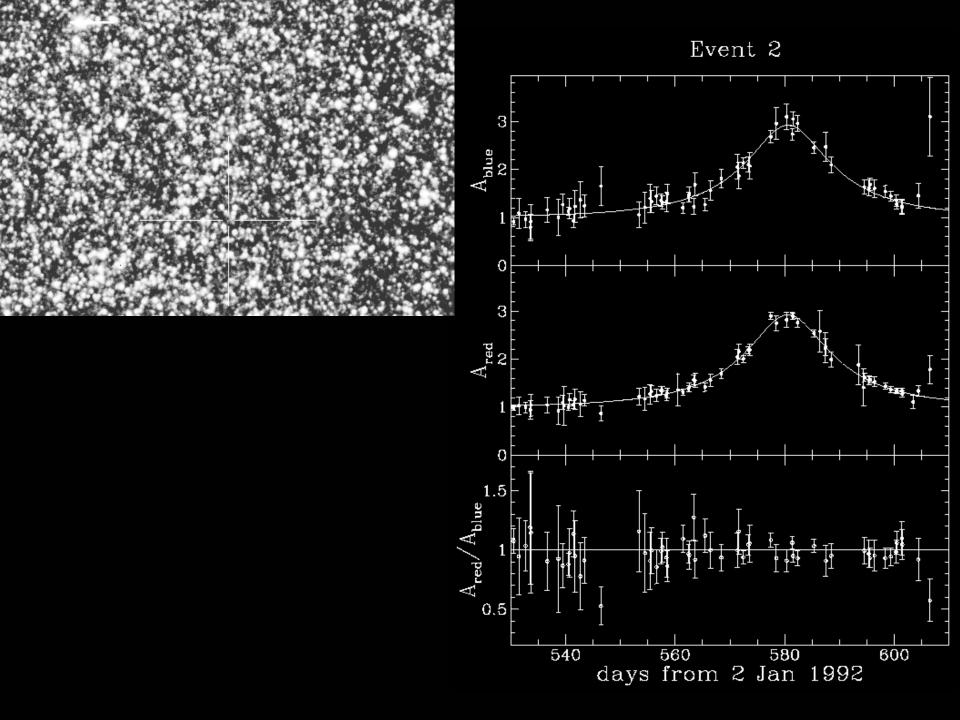
- Modified Newtonian dynamics
- Planets
- Mass disadvantaged stars
   brown red white
- Black holes

gravitational microlensing









# Most of the universe is Cark!

- Modified Newtonian dynamics
- Planets
- Mass disadvantaged stars

brown red white

- Black holes
- Fossil remnant of the big bang



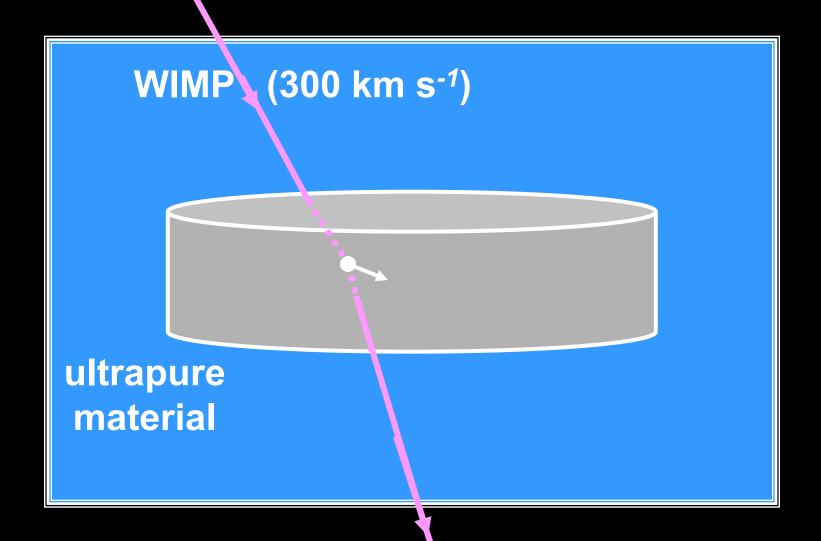
# A WIMPY IDEA



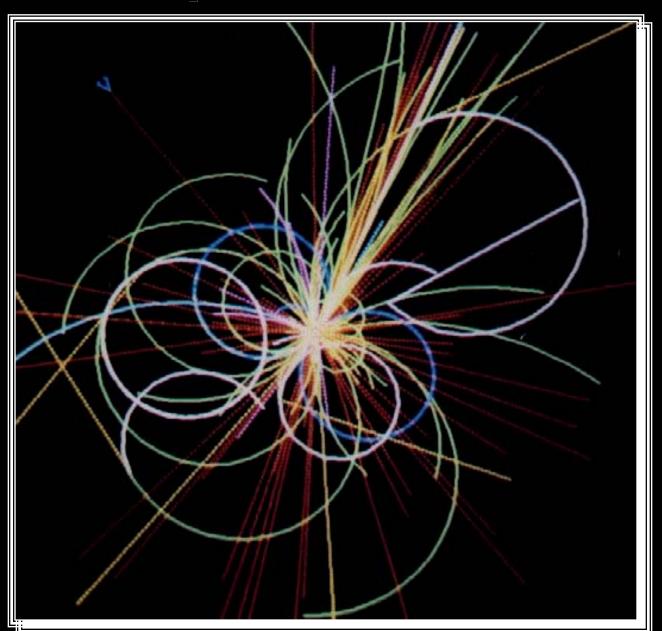
- Most of the universe is invisible
- Dominated by the rest mass of an elementary particle
  - present in the primordial soup
  - massive
  - neutral
  - weakly interacting
  - slow
  - stable



### Direct detection



# Make wimps in the laboratory

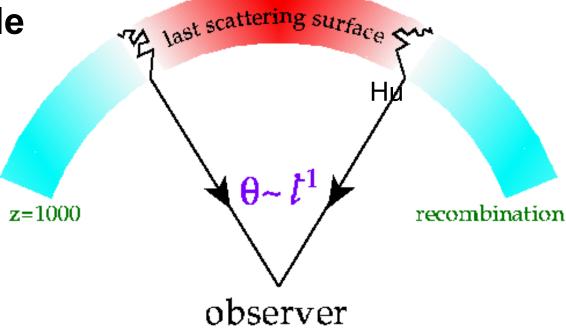


### Primordial soup

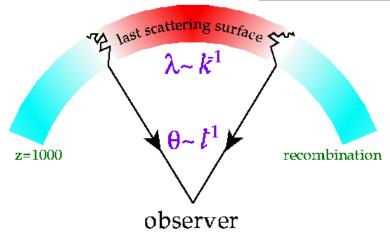
### **KNOWN INGREDIENTS:** 56% QUARKS 16% GLUONS (STRONG FORCE) 9% ELECTRON-LIKE PARTICLES (WEAK FORCE) 9% W's AND Z's 5% NEUTRINOS 2% PHOTONS (ELECTROMAGNETIC FORCE) 2% GRAVITONS (GRAVITATIONAL FORCE) 1% HIGGS BOSONS (???? DARK MATTER

### Acoustic peaks

- At recombination, baryon-photon fluid undergoes "acoustic oscillations"
- Compressions and rarefactions change  $T_{\gamma}$
- Peaks in  $\Delta T_{\gamma}$  correspond to extrema of compressions and rarefactions
- Multipole number corresponds to a physical length scale



### Acoustic peaks



# Observer $\frac{\Omega_0 \quad \Omega_{\Lambda}}{\Omega_0 \quad \Omega_{\Lambda}}$ beam 1.0 0.0 0.3 0.0 0.3 0.7 cosmic variance

 $\Omega_{\text{TOTAL}} = 1 \pm 0.1$ 

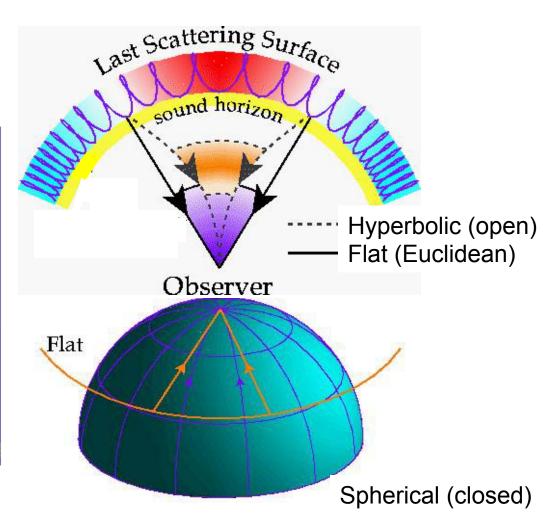
 $10^{2}$ 

 $10^{3}$ 

10

 $10^{2}$ 

# Sound travel distance known Observed $I_{peak}$ ~ geometry



# Dark energy?

Space and time are related.

1905

Space is dynamical (curved, warped, bent).

1915

Empty space has a weight.

1917





the weight of space

Cosmological

constant

(Dark energy)

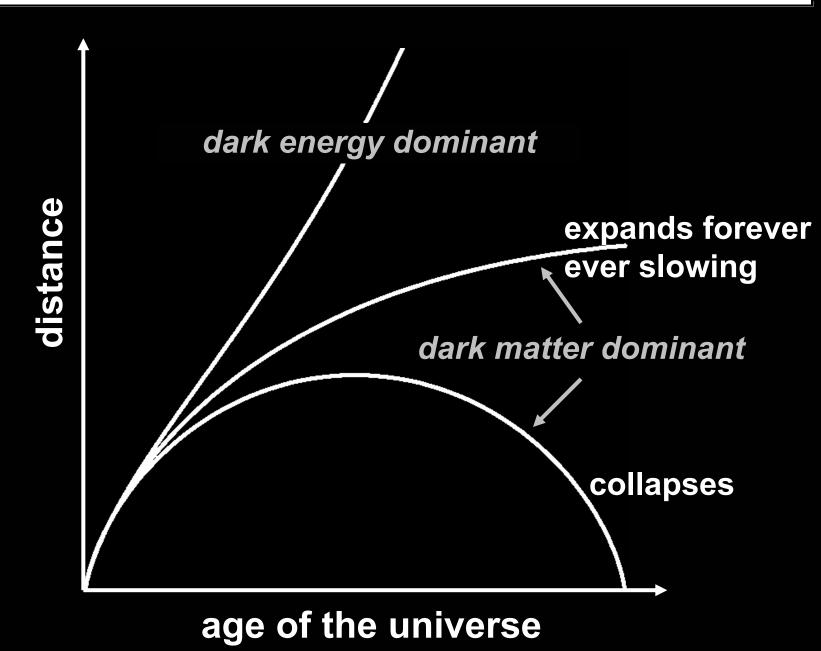
1917 Einstein proposed cosmological constant.

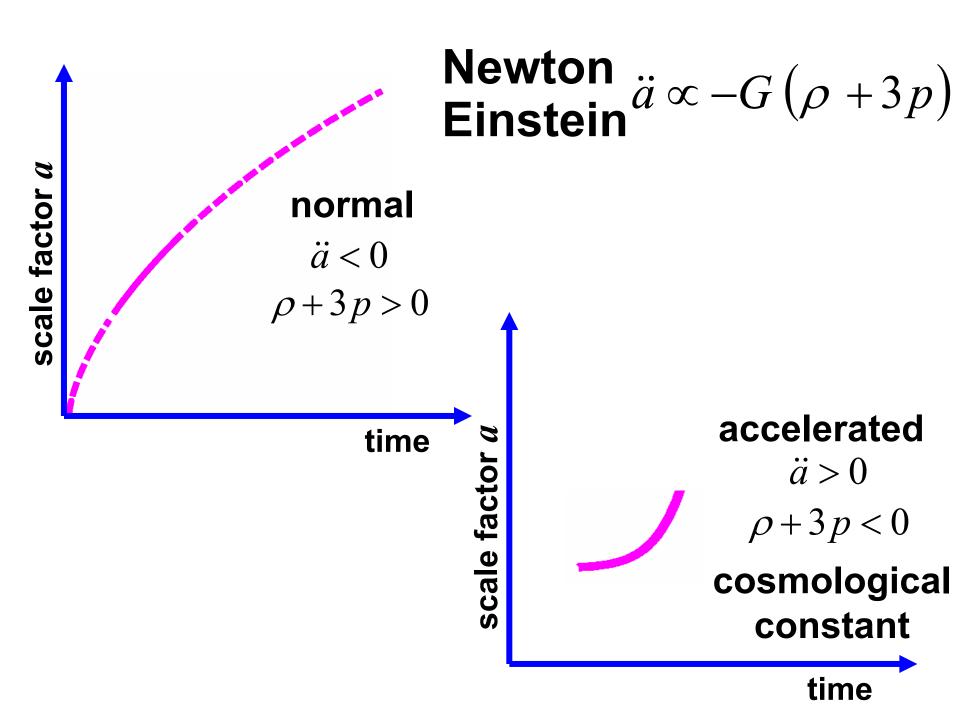
1929 Hubble discovered expansion of the Universe.

1934 Einstein called it "my biggest blunder."

1998 Astronomers found evidence for it.

### Cosmological constant (dark energy)





# Vacuum pressure



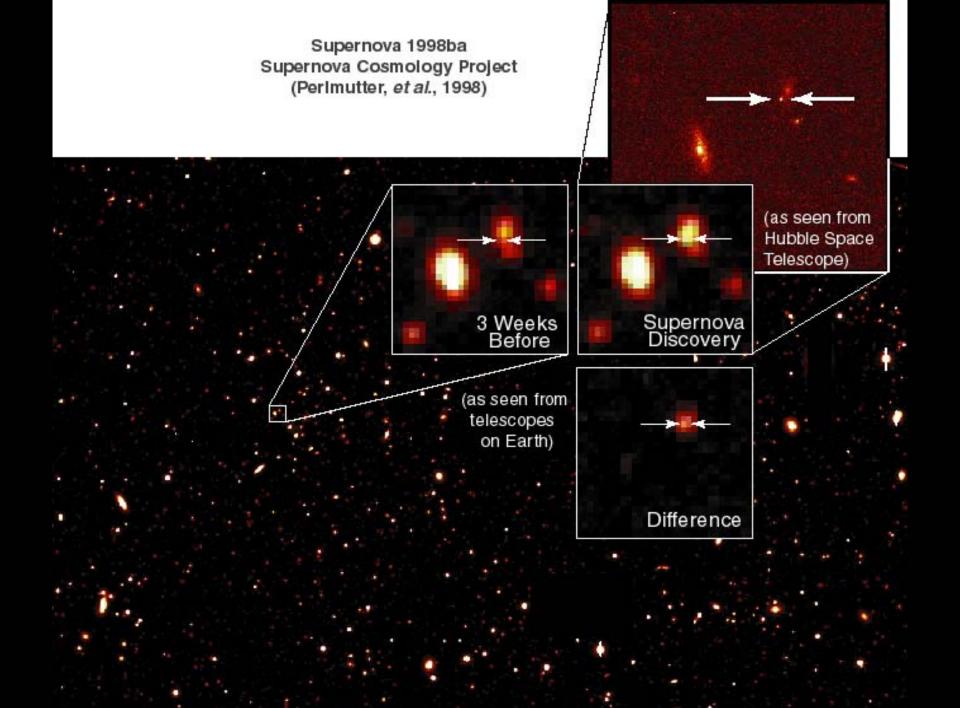


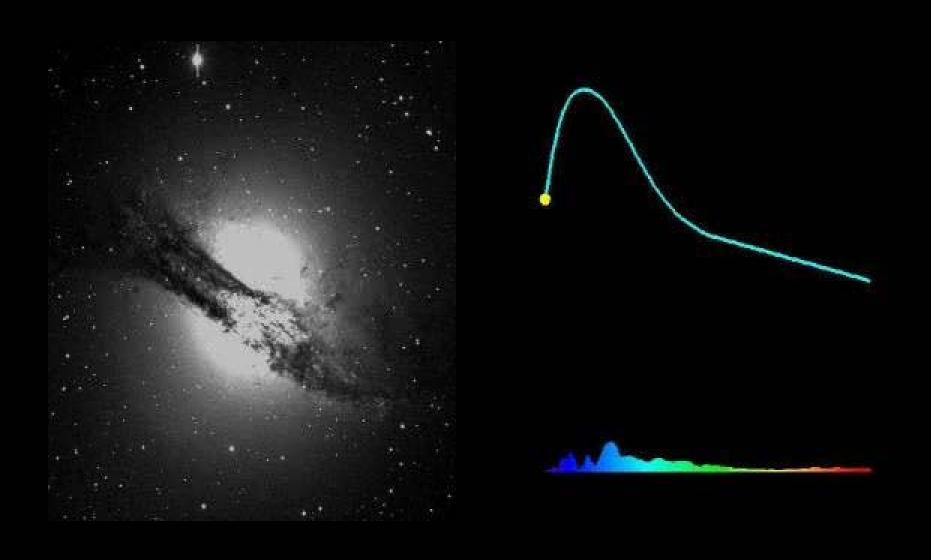
 $E_2 > E_1$ 

had to pull piston "negative pressure"

### SN 1987A in the LMC

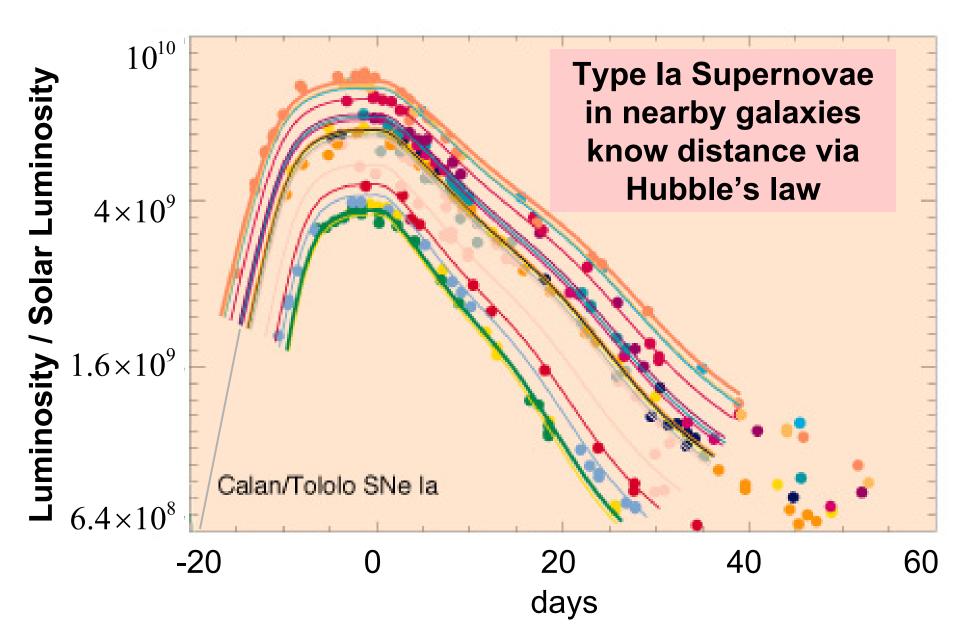




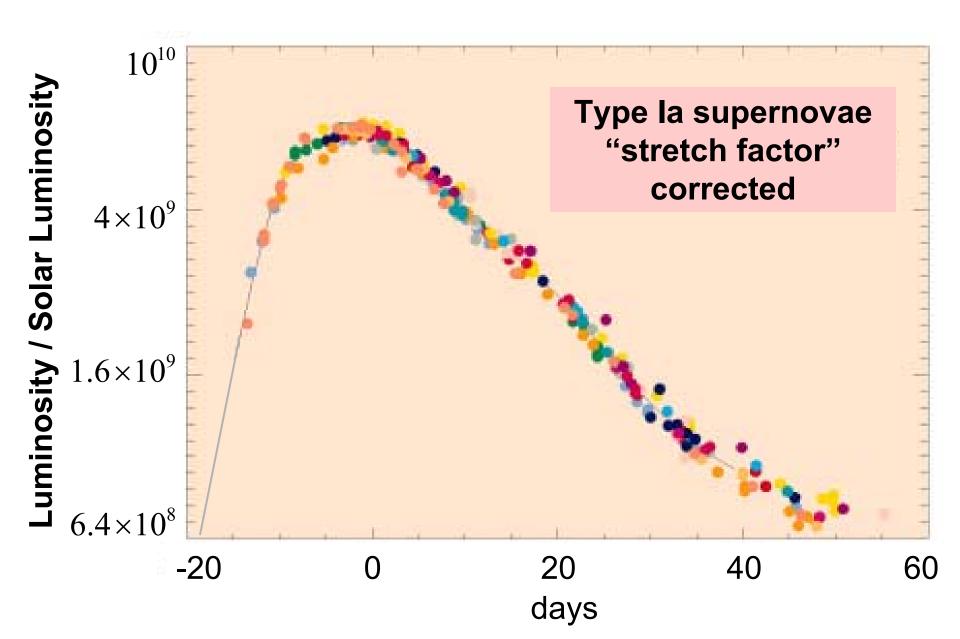


Centaurus A

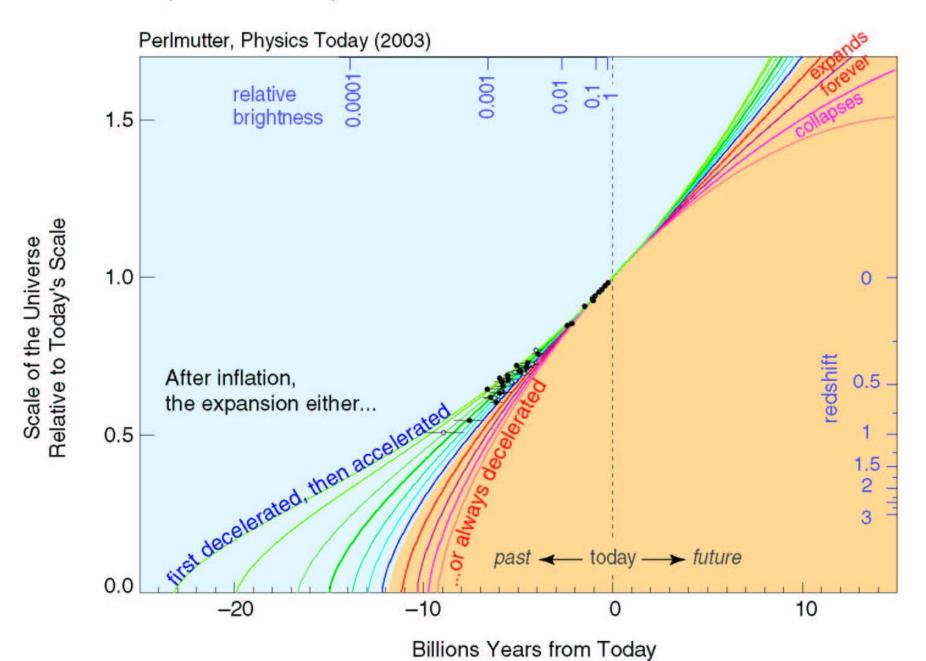
### Type la supernova are standard candles



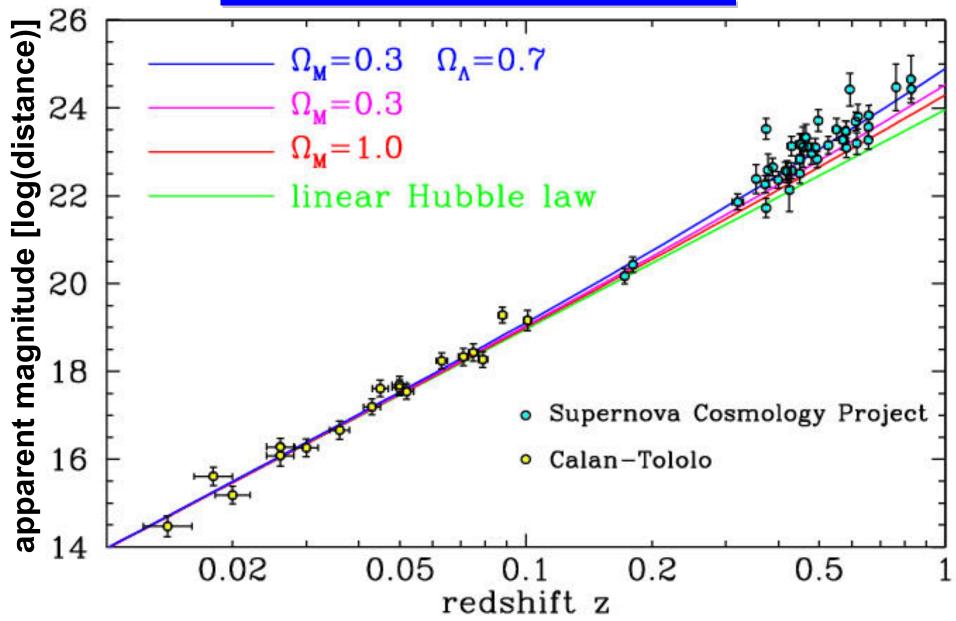
### Type la supernova are standard candles



### **Expansion History of the Universe**



Type la supernova



# Cosmological constant (Dark energy)

Mass density of space:  $10^{-30}$  g cm<sup>-3</sup>

$$10^{-30} \, \mathrm{g \ cm^{-3}}$$

The unbearable lightness of nothing!

### Cosmo-illogical constant?